

## NOTES FOR 28 JAN (TUESDAY)

### 1. RECAP

- (1) Recalled definitions and examples of Vector bundles. Did some constructions of new vector bundles from old ones.
- (2) Defined metrics, and gave examples.

### 2. RIEMANNIAN MANIFOLDS AND METRICS ON VECTOR BUNDLES

Recall the definition of an induced metric

**Definition 2.1.** If  $g$  is a metric on  $M$  and  $S \subset M$  is an embedded submanifold, then  $g$  induces a metric  $g|_S$  on  $S$  given by  $g_p|_S(v_S, w_S) = g_p(i_*v_S, i_*w_S)$ .

- (1)  $S^2 \subset \mathbb{R}^3$ . First write the metric in  $\mathbb{R}^3$  in spherical coordinates  $z = r \cos(\theta)$ ,  $x = r \sin(\theta) \cos(\phi)$ ,  $y = r \sin(\theta) \sin(\phi)$ . Thus,  $g_{Euc} = dr \otimes dr + r^2 d\theta \otimes d\theta + r^2 \sin^2(\theta) d\phi \otimes d\phi$ . Now when we restrict to the unit sphere, the tangent vectors do not include  $\frac{\partial}{\partial r}$ . Thus,  $g_{Sphere} = d\theta \otimes d\theta + \sin^2(\theta) d\phi \otimes d\phi$
- (2) Suppose  $z = f(x, y)$  is the graph of a function, then  $g_{Induced} = dx \otimes dx + dy \otimes dy + (\frac{\partial f}{\partial x})^2 dx \otimes dx + (\frac{\partial f}{\partial y})^2 dy \otimes dy + \frac{\partial f}{\partial x} \frac{\partial f}{\partial y} (dx \otimes dy + dy \otimes dx)$ .

Now we write down the volume forms of most of the above examples :

- (1)  $vol_{Euc} = dx^1 \wedge dx^2 \wedge \dots \wedge dx^n$ .
- (2) In polar coordinates in  $\mathbb{R}^2$ ,  $vol_{Euc} = \sqrt{\det(g)} dr \wedge d\theta = r dr \wedge d\theta$ .
- (3) For the circle,  $vol = d\theta$ .

Suppose  $\gamma : [0, 1] \rightarrow M$  is a smooth path. Then, define its length as  $L(\gamma) = \int_0^1 \sqrt{g(\frac{d\gamma}{dt}, \frac{d\gamma}{dt})} dt$ . A piecewise  $C^1$ , regular (meaning that  $\gamma' \neq 0$  throughout) curve satisfying the following equation is called a geodesic.

$$(2.1) \quad \begin{aligned} \frac{d^2 \gamma^r}{dt^2} + \Gamma_{ij}^r \frac{d\gamma^i}{dt} \frac{d\gamma^j}{dt} &= 0 \\ \Gamma_{ij}^r &= g^{rl} \frac{1}{2} \left( \frac{\partial g_{il}}{\partial x^j} + \frac{\partial g_{jl}}{\partial x^i} - \frac{\partial g_{ij}}{\partial x^l} \right) \end{aligned}$$

It can be proved that every geodesic is actually smooth, can be parametrised by its arc-length, and that arc-length parametrised geodesics are precisely the critical points of the length functional. Moreover, the function  $d(p, q) = \inf L(p, q)$  over all the piecewise  $C^1$  paths joining  $p$  and  $q$  is a metric and that the topology induced by it is the same as the original topology of the manifold.

Now we note that geodesics exist locally, and that if  $\gamma$  is a geodesic, then so is  $\gamma(ct)$ . In fact, we have the following result.

**Theorem 2.2.** Let  $p \in M$ . Then there is a neighbourhood  $U_o$  of  $p$  and a number  $\epsilon_p > 0$  such that for every  $q \in U$  and every tangent vector  $v \in T_q M$  with  $\|v\| < \epsilon_p$  there is a unique geodesic  $\gamma_v : (-2, 2) \rightarrow M$  satisfying  $\gamma_v(0) = q$ ,  $\frac{d\gamma_v}{dt}(0) = v$ .

If  $v \in T_q M$  is a vector for which there is a geodesic,  $\gamma : [0, 1] \rightarrow M$  satisfying  $\gamma(0) = q$  and  $\gamma'(0) = v$  then we define  $\exp_q(v) = \gamma_v(1)$ . The geodesic itself can be described as  $\gamma(t) = \exp_q(tv)$  (by the uniqueness theorem for ODE). By the smooth dependence on parameters of an ODE,  $\exp_q(v)$  depends smoothly on  $q$  and on  $v$  and defines a smooth map  $\exp_q : T_q M \rightarrow M$ .

Note that  $(\exp_q)_{v*} : T_v(T_q M) \simeq T_q M \rightarrow T_{\exp_q(v)} M$  is its pushforward. We claim that

**Theorem 2.3.**  $(\exp_q)_{0*} = Id$  and hence  $\exp_q$  is a local diffeomorphism around  $\vec{0}$ .

*Proof.* Clearly the first statement and the inverse function theorem imply the second. Now if  $v \in T_q M$ , we need to obtain a curve  $c(t) \in T_q M$  such that  $c(0) = 0$ ,  $c'(0) = v$ , and  $\frac{d\exp_q(c(t))}{dt}|_{t=0} = v$ . Let  $c(t) = tv$ . Then  $\exp_q(c(t)) = \exp_q(tv)$  which is the time- $t$  geodesic starting at  $q$  pointing along  $v$  at  $t = 0$ . Thus we are done.  $\square$

In fact, we can say more.

**Theorem 2.4.** *Geodesics are locally length minimising. Moreover, if  $p \in M$ , there exists a geodesic ball  $B_{\epsilon_p}(p)$  such that every two points in the ball can be connected by a unique length minimising geodesic lying in the ball and such that the exponential map is a diffeomorphism restricted to the ball. Such a ball is called a geodesically convex ball.*

Now we make a definition of a useful coordinate system.

**Definition 2.5.** Given  $q \in M$ , the coordinate system defined by  $\exp_q : U \subset T_q M \rightarrow M$  is called a geodesic normal coordinate system at  $q$  (after choosing coordinates on  $U$  that is).

This set of coordinates is extremely useful. In fact,

**Theorem 2.6.** *There is a geodesic normal coordinate system  $v$  at  $p$ ,  $g_{ij}(p) = \delta_{ij}$  and  $\frac{\partial g_{ij}}{\partial v^k}(p) = 0$ .*

*Proof.* Choose coordinates  $x^\mu$  so that  $g_{\mu\nu}(p) = \delta_{\mu\nu}$ . (This can be easily accomplished by taking any coordinate system and rotating it so as to diagonalise  $g$ .) Let  $v^i$  be coordinates in  $T_p M$ . Now  $\exp$  is a local diffeomorphism. So  $x^\mu(v^j) = x^\mu \circ \exp(v^j)$  is a change of coordinates in a small neighbourhood.

Note that since  $\exp_{0*} = Id$ ,  $\frac{\partial x^\mu}{\partial v^j}|_{v=0} = \delta_j^\mu$ . Now  $\tilde{g}_{ij} = g_{\mu\nu} \frac{\partial x^\mu}{\partial v^i} \frac{\partial x^\nu}{\partial v^j}$ . So it is easy to see that  $\tilde{g}_{ij}(0) = \delta_{ij}$ . Since the geodesics through  $p$  are linear in this coordinate system, we see that the Christoffel symbols  $\tilde{\Gamma}_{ij}^r(0) = 0$ . It is easy to see that if the Christoffel symbols are 0, then so are all first partial derivatives of the metric.  $\square$

More generally, any coordinate system in which the metric at  $p$  is standard upto first order is called a normal coordinate system at  $p$ .

Actually, we can prove the existence of normal coordinates in much simpler manner even without reference to geodesics.

**Theorem 2.7.** *There is a normal coordinate system  $y$  at  $p$ .*

*Proof.* Choose any coordinate system at  $x$  at  $p$  such that  $x = 0$  is  $p$ . Using a linear map, we may diagonalise  $g$  at  $p$ . So without loss of generality,  $\tilde{g}_{\mu\nu} = \delta_{\mu\nu} + a_{\mu\nu\alpha} x^\alpha + O(x^2)$ . (Note that  $a_{\mu\nu\alpha} = a_{\nu\mu\alpha}$ .) Change the coordinates to  $y$  such that  $x(y)^i = y^i + b_{jk}^i y^j y^k$  where  $b_{jk}^i = b_{kj}^i$ . Now

$$\begin{aligned} g_{ij} &= \tilde{g}_{\mu\nu} \frac{\partial x^\mu}{\partial y^i} \frac{\partial x^\nu}{\partial y^j} = (\delta_{\mu\nu} + a_{\mu\nu\alpha} y^\alpha + O(y^2)) (\delta_i^\mu + b_{ik}^\mu y^k) (\delta_j^\nu + b_{jk}^\nu y^k) \\ (2.2) \quad &= \delta_{ij} + a_{ijk} y^k + (b_{ijk} + b_{jik}) y^k + O(y^2) \end{aligned}$$

So we just need to choose  $b$  so that  $a_{ijk} = -b_{ijk} - b_{jik} \forall k$ . So take  $b = -\frac{a}{2}$ .  $\square$

It is natural to ask if there is a geodesic normal coordinate system to the second order. Shockingly enough, there isn't (in general). In fact,

**Theorem 2.8.** *There exists a  $(0,4)$  tensor (called the Riemann curvature tensor of  $g$ ) which is locally  $R_{\mu\nu\alpha\beta}$  such that in geodesic normal coordinates,*

$$(2.3) \quad g_{\mu\nu} = \delta_{\mu\nu} - \frac{1}{3}R_{\mu\alpha\nu\beta}(0)x^\mu x^\nu + O(x^3)$$

where in these coordinates,  $R_{ijkl}(0) = \frac{1}{2} \frac{\partial^2 g_{jk}}{\partial x^i \partial x^l}(0) + \frac{1}{2} \frac{\partial^2 g_{il}}{\partial x^j \partial x^k}(0) - \frac{1}{2} \frac{\partial^2 g_{jl}}{\partial x^i \partial x^k}(0) - \frac{1}{2} \frac{\partial^2 g_{ik}}{\partial x^j \partial x^l}(0)$ . In fact, all the other terms in the Taylor expansion depend only on  $R$  and its derivatives. So there is a change of coordinates such that  $g$  is Euclidean everywhere, then, since the Euclidean coordinates are geodesically normal, the Riemann curvature tensor is identically 0.

So one can prove that one cannot draw a map of any part of Bangalore on a piece of paper such that distances are to scale, by calculating the curvature of the sphere with the metric induced from the Euclidean space. It turns out to be a non-zero tensor. We will return to curvature later on in a different way. This theorem is to show you that the notion of curvature is "forced" upon us. (It is not an artificial definition.)